# Chiral matter wavefunctions in warped compactifications

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# Outline

- 1. Motivation and setup
- 2.Unmagnetized branes
- 3.Magnetized branes
- 4. Warped kinetic terms
- 5.Conclusions

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## Motivation

•Warping plays an important role in string models:

- •Generation of the electroweak hiearchy [RS; GKP]
- Stabilization of moduli [KKLT;...]
- •Late-time acceleration [KKLT;...] and inflation [KKLMMT;...]
- •Gauge/gravity duality
- •In type II theories, realistic models require open strings
- •A 4D effective action is valuable for detailed phenomenology

# Motivation (cont.)

- Alternatively, consider an F-theory compactification
- In such constructions must satisfy the D3-tadpole condition

$$\frac{\chi(X)}{24} = N_{\rm D3} + \int_{B_3} H^{(3)} \wedge F^{(3)}$$

 Additional ingredients will cause warping which will modify 4D EFT



#### Warped effective field theory

- In a flat background, CFT techniques can be used to determine an EFT (Kähler metrics, Yukawa couplings [Lüst et. al.; Cvetič et. al.; Ibáñez et. al;...])
- However, warping in type II usually involves Ramond-Ramond fluxes and it is difficult to make use of CFT techniques



### Warped effective field theory

- Alternative: Dimensional reduction of a higher dimensional EFT
- Requires knowledge of wavefunctions of light <sup>-</sup> degrees of freedom
- Here, my focus is on open string modes (see
   B. Underwood's talk for closed strings)

$$-\frac{1}{2g_8^2} \int_{\mathcal{W}} \mathrm{d}^8 x \sqrt{g} F^2$$

$$A_\mu(x,y) = A_\mu(x) s(y)$$

$$-\frac{1}{4g_8^2} \int_{\mathbb{R}^{1,3}} \mathrm{d}^4 x F_{\mu\nu} F^{\mu\nu} \int_{\mathcal{S}_4} \mathrm{d}^4 y \sqrt{g} s^2$$

$$\boxed{\frac{1}{g_4^2} = \frac{\mathcal{V}_{\mathcal{S}_4}}{g_8^2}}$$

## Setup

- •For simplicity, compactify IIB on  $\mathbb{T}^6 = \mathbb{T}_1^2 \times \mathbb{T}_2^2 \times \mathbb{T}_3^2$  $ds_{10}^2 = e^{2\alpha} dx_4^2 + e^{-2\alpha} dz^m d\overline{z}^{\overline{m}}$   $F^{(5)} = (1 + *_{10}) F_{\text{ext}}^{(5)} \quad F_{\text{ext}}^{(5)} = e^{4\alpha} \wedge \text{dvol}_{\mathbb{R}^{1,3}} \quad G^{(3)} = 0$
- •Add two probe D7 branes

$$D7_1: z^3 = M_3 z^2$$

$$D7_2: z^3 = -M_3 z^2$$

• Gives a U (1)  $\times$  U (1) gauge symmetry



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## Single D-brane

•For a single D7-brane, the 4D bosonic d.o.f. are

- •Gauge boson  $A_{\mu}$
- •Wilson lines  $A_a$
- ullet Worldvolume deformations  $\Phi$
- Dynamics described by the DBI+CS action

10

## Intersections as Higgsing

•When the branes are coincident, the symmetry is enhanced to U(2). The transverse fluctuations are promoted to an adjoint-valued scalar  $\Phi$ 

Position of D7<sub>1</sub> Position of D7<sub>2</sub>  

$$\Phi = \begin{pmatrix} \phi^a & \phi^- \\ \phi^+ & \phi^b \end{pmatrix}$$
7<sub>1</sub>-7<sub>2</sub> strings (bifundamental)

•vevs for  $\phi^{a,b}$  correspond to background D7 positions

$$\langle \phi^a \rangle = \lambda^{-1} M_3 z^2 \quad \left\langle \phi^b \right\rangle = -\lambda^{-1} M_3 z^2$$



#### Myers action

 Bosonic fluctuations governed by the Myers action  $S_{\mathrm{D7}} = S_{\mathrm{D7}}^{\mathrm{DBI}} + S_{\mathrm{D7}}^{\mathrm{CS}} \operatorname{symmetrization} \\ S_{\mathrm{D7}}^{\mathrm{DBI}} = -\tau_{\mathrm{D7}} \int_{\mathcal{W}} \mathrm{d}^{8} x \operatorname{Str} \left\{ \left( \operatorname{Im} \tau \right)^{-1} \sqrt{\det M_{\alpha\beta} \det Q_{j}^{i}} \right\}$ interior product  $S_{\mathrm{D7}}^{\mathrm{CS}} = \tau_{\mathrm{D7}} \int_{\mathcal{W}} \mathrm{Str} \left\{ \Pr \left[ \mathrm{e}^{\mathrm{i}\lambda\iota_{\Phi}\iota_{\Phi}} \mathcal{C} \wedge \mathrm{e}^{B^{(2)}} \right] \wedge \mathrm{e}^{\lambda F^{(2)}} \right\}$ non-Abelian pullback:  $P[v_{\alpha}] = v_{\alpha} + \lambda v_i D_{\alpha} \Phi^i$ where:  $M_{\alpha\beta} = P \left[ E_{\alpha\beta} + \left( \operatorname{Im} \tau \right)^{-1/2} E_{\alpha i} \left( Q^{-1} - \delta \right)^{ij} E_{j\beta} \right] + \lambda \left( \operatorname{Im} \tau \right)^{1/2} F_{\alpha\beta}$  $E_{MN} = g_{MN} + \left(\operatorname{Im} \tau\right)^{1/2} B_{MN} \qquad Q_i^i = \delta_i^i - i\lambda \left[\Phi^i, \Phi^k\right] \left(\operatorname{Im} \tau\right)^{-1/2} E_{ki}$ 

## Myers action (cont.)

Bulk fields given as a non-Abelian Taylor expansion

adjoint valued  

$$\Psi[\Phi] = \sum_{n=0}^{\infty} \frac{\lambda^n}{n!} \Phi^{i_1} \cdots \Phi^{i_n} \partial_{i_1} \cdots \partial_{i_n} \Psi_0$$
neutral  

$$\Psi_0 + \mathcal{O}(\lambda)$$
need small angle

- •Leading order in  $\alpha'$ , action is Higgsed warped Yang-Mills
- Equations of motion are second order and hard to solve in general

#### Fermionic action

•To get first order equations, can use fermionic action

•d.o.f. encoded in two 10D M-W spinors 
$$\Theta = \begin{pmatrix} \theta_1 \\ \theta_2 \end{pmatrix}$$

•Abelian case: [Martucci, Rosseel, Van denBleeken, Van Proeyen;...]

$$S_{\mathrm{D7}}^{\mathrm{F}} = \frac{1}{g_{8}^{2}} \int \mathrm{d}^{8} x \left(\mathrm{Im}\,\tau\right)^{-1} \sqrt{\det M_{\alpha\beta}} \bar{\Theta} P_{-}^{\mathrm{D7}} \left[ \left(\mathcal{M}^{-1}\right)^{\alpha\beta} \Gamma_{\beta} \mathcal{D}_{\alpha} - \mathcal{O} \right] \Theta$$

$$\int \int \delta_{\epsilon} \Psi_{M} = \mathcal{D}_{M} \epsilon$$
Involves  $\Gamma_{\mathrm{D7}}$ 

$$\delta_{\epsilon} \lambda = \mathcal{O} \epsilon$$

#### Fermionic action (cont.)

•In a warped geometry ( $G^{(3)} = 0$ ),

•After  $\kappa$ -fixing

$$S_{\mathrm{D7}}^{\mathrm{F}} = \frac{1}{g_{8}^{2}} \int_{\mathcal{W}} \mathrm{d}^{8} x \,\bar{\theta} \bigg\{ \mathrm{e}^{-\alpha} \partial_{\mathbb{R}^{1,3}} + \mathrm{e}^{\alpha} \partial_{\mathbb{T}^{4}} + \mathrm{e}^{\alpha} \frac{1}{2} \partial_{\mathbb{T}^{4}} \alpha \big( 1 + 2\Gamma_{\mathcal{S}^{4}} \big) \bigg\} \theta$$

$$4\text{-cycle chirality}$$

•Non-Abelian modification (leading  $\alpha'$ ) [Wynants]:

$$\partial \to D \qquad \delta \mathcal{L} = -\mathrm{i}\bar{\theta}\mathrm{e}^{-\alpha}\Gamma_i [\Phi^i, \theta]$$

and take trace

# Adjoint fields

 Warping effect on adjoint zero-mode wavefunctions are mild

Vector multiplet: $A_{\mu} \sim \text{const}$  $\psi_0 \sim e^{-3\alpha/2}$ Wilson line multiplet: $A_m \sim \text{const}$  $\psi_{1,2} \sim e^{\alpha/2}$ Deformation multiplet: $\phi \sim \text{const}$  $\psi_3 \sim e^{-3\alpha/2}$ 

Consistent with supersymmetry

#### Equations of motion

•For the bifundamental modes, take the ansatz

$$\psi_{0,3}^{\mp} = e^{-3\alpha/2} \chi_{0,3}^{\mp}$$
  
 $\chi$   
gaugino, modulino

$$\psi_{1,2}^{\mp} = e^{\alpha/2} \chi_{1,2}^{\mp}$$

$$\uparrow$$
wilsonini

•Equations of motion:

 $0 = \partial_1 \chi_1^{\mp} + \partial_2 \chi_2^{\mp} + e^{-4\alpha} D_3^{\mp} \chi_3^{\mp}$   $0 = \partial_1 \chi_0^{\mp} + \partial_2^* \chi_3^{\mp} - D_3^{\pm *} \chi_2^{\mp}$   $0 = \partial_1^* \chi_3^{\mp} - \partial_2 \chi_0^{\mp} - D_3^{\pm *} \chi_1^{\mp}$  $0 = \partial_1^* \chi_2^{\mp} - \partial_2^* \chi_1^{\mp} + e^{-4\alpha} \hat{D}_3^{\mp} \chi_0^{\mp}$ 

$$D_3^{\mp} = \mp \mathrm{i} M_3 \bar{z}^{\bar{2}}$$

#### **BPS** conditions

- •For a single D7 brane, the equations of motion follow from F- and D-flatness conditions: [Jockers, Louis; Martucci]  $\begin{array}{ll} \text{fundamental} & W = \int_{\mathcal{S}_4} \mathbf{P}[\gamma] \wedge \mathbf{e}^{\lambda F^{(2)}} & D = \int_{\mathcal{S}_4} \mathbf{P}[\operatorname{Im} \eta] \wedge \mathbf{e}^{\lambda F^{(2)}} & \text{warped} \\ \hline \mathbf{3}\text{-form} & \mathbf{d}\gamma = \Omega \wedge \mathbf{e}^{B^{(2)}} & \eta = \mathbf{e}^{2\alpha} \operatorname{Im} \tau \, \mathbf{e}^{\mathbf{i}J} \wedge \mathbf{e}^{B^{(2)}} & \text{form} \end{array}$ 
  - - Comparing to the CS-action, the non-Abelian version should be (see also [Butti et. al.])

$$W = \int_{\mathcal{S}_4} \operatorname{Str} \left\{ \operatorname{P} \left[ \operatorname{e}^{\mathrm{i}\lambda\iota_{\Phi}\iota_{\Phi}} \gamma \right] \wedge \operatorname{e}^{\lambda F^{(2)}} \right\} \ D = \int_{\mathcal{S}_4} \operatorname{S} \left\{ \operatorname{P} \left[ \operatorname{e}^{\mathrm{i}\lambda\iota_{\Phi}\iota_{\Phi}} \operatorname{Im} \eta \right] \wedge \operatorname{e}^{\lambda F^{(2)}} \right\}$$

These yield the previous equations of motion with  $\psi_0 = 0$ 

## Unwarped zero mode

 In the absence of warping, the zero modes are exponentially localized on the intersection



 Mixture of deformation modulus and Wilson line of the un-Higgsed theory

#### Warped zero mode

- •For arbitrary warping, no simple analytic solution
- In the weak warping case, can treat the warping as a perturbation

$$e^{-4\alpha} = 1 + \epsilon\beta \qquad \epsilon \ll 1$$

•Can then expand the warped zero mode in terms of the unwarped massive modes

#### Unwarped spectrum

•The equation of motion for the massive modes is

$$\mathbf{D}^{\mp} \mathbf{X}_{\lambda}^{\mp} = m_{\lambda} \mathbf{X}_{\lambda}^{\pm *} \quad \mathbf{D}^{\mp} = \begin{pmatrix} 0 & \partial_{1} & \partial_{2} & D_{3}^{\mp} \\ -\partial_{1} & 0 & D_{3}^{\pm *} & -\partial_{2}^{*} \\ -\partial_{2} & -D_{3}^{\pm *} & 0 & \partial_{1}^{*} \\ -D_{3}^{\mp} & \partial_{2}^{*} & -\partial_{1}^{*} & 0 \end{pmatrix} \quad \mathbf{X}_{\lambda}^{\mp} = \begin{pmatrix} \chi_{0}^{\mp} \\ \chi_{1}^{\mp} \\ \chi_{2}^{\mp} \\ \chi_{3}^{\mp} \end{pmatrix}$$

• Easiest to work in a rotated basis  $\mathbf{X}'^{\mp} = \mathbf{J}^{-1}\mathbf{X}^{\mp}$ 

$$\mathbf{J} = \begin{pmatrix} 1 & & \\ & 1 & & \\ & & 1/\sqrt{2} & i/\sqrt{2} \\ & & i/\sqrt{2} & 1/\sqrt{2} \end{pmatrix} \qquad \begin{array}{l} \partial_1 \to \partial_1 \\ \partial_2 \to \hat{D}_2^{\prime \mp} = \frac{1}{\sqrt{2}} \left( \partial_2 \pm M_3 \bar{z}^{\bar{2}} \right) \\ \partial_3 \to \hat{D}_3^{\prime \mp} = \frac{1}{\sqrt{2}} \left( \partial_2 \mp M_3 \bar{z}^{\bar{2}} \right) \end{array}$$

#### Unwarped spectrum (cont.)

- •Boundary conditions:
  - Periodicity along  $\mathbb{T}_1^2$
  - Localized on intersection
- - sector modes built from ladder operators (giving two simple harmonic oscillator algebras) and Fourier modes



#### Unwarped spectrum (cont.)

•Unwarped spectrum:

$$m_{\lambda}^{2} = m^{2} + n^{2} + M_{3}(l + p + 1)$$

$$m_{\lambda}^{2} = m^{2} + n^{2} + M_{3}(l + p + 1)$$

$$\Phi_{\lambda}^{\prime-} = \left(\varphi_{mnlp}^{-}, 0, 0, 0\right)$$

$$\Phi_{\lambda}^{\prime-}=\left(0,\varphi_{mnlp}^{-},0,0\right)$$

$$m_{\lambda}^2 = m^2 + n^2 + M_3(l+p)$$

$$m_{\lambda}^{2} = m^{2} + n^{2} + M_{3}(l + p + 2)$$

$$\Phi_{\lambda}^{\prime-} = \left(0, 0, \varphi_{mnlp}^{-}, 0\right)$$

$$\Phi_{\lambda}^{\prime-} = \left(0, 0, 0, \varphi_{mnlp}^{-}\right)$$

#### Expanding the warped zero mode

•Write the warped zero mode as

$$\mathbf{X}^{-} = \boldsymbol{\Phi}_{0}^{-} + \sum_{\lambda} c_{\lambda} \boldsymbol{\Phi}_{\lambda}^{-}$$
  
unwarped modes

•To leading order

$$c_{\lambda} = \frac{1}{m_{\lambda}^2} \int_{\mathcal{S}_4} \mathrm{d}^4 y \left( \mathbf{\Phi}_{\lambda}^- \right)^* \cdot \left( \mathbf{D}_0^+ \right)^* \beta \mathbf{K}^- \mathbf{\Phi}_0^-$$

where

$$\mathbf{D}^{-} = \mathbf{D}_{0}^{-} + \epsilon \beta \mathbf{K}^{-} + \mathcal{O}(\epsilon^{2})$$

• Examples given in [1011.xxx]

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# Chirality

- •Without magnetic flux, the spectrum is vector-like
- In order to have a chiral theory, the intersection must be magnetized

$$\frac{1}{2\pi} \int_{\mathbb{T}^2} F^{(2)} = M_1 \sigma_3$$

•SUSY requires [Marino, Minasian, Moore, Strominger; ...]

$$F^{(2)} = - *_4 F^{(2)}$$

$$\bigwedge_{\text{Hodge-*on} S_4}$$

#### Unwarped zero modes

•For example, if  $M_1 > 0$ , only the –-sector has zero modes

• Due to magnetic flux, wavefunction are quasi-periodic [Cremades, Ibáñez, Marchesano;...]

$$\varphi_{0}^{j,-} = e^{-\kappa |z^{2}|^{2}} e^{2\pi i M_{1} z^{1} \operatorname{Im} z^{1}} \vartheta \begin{bmatrix} j/2M_{1} \\ \uparrow & 0 \end{bmatrix} (2M_{1} z^{1}, i2M_{1})$$
  
$$\kappa = \sqrt{\left(\frac{M_{1}}{2}\right)^{2} + M_{3}^{2}} \qquad j = 0, \dots, 2M_{1} - 1$$

families orthogonal:  $\int_{\mathcal{S}_4} \mathrm{d}^4 y \left(\varphi_0^{j,-}\right)^* \varphi_0^{k,-} = \delta^{kj}$ 

#### Warped zero modes

- As in unmagnetized case, warped zero mode has no general simple analytic solution
- •Again, expand in unwarped massive modes
- Spectrum built from three QSHO algebras

$$\varphi_{nlp}^{j,-} = (iD_1'^-)^n [i(D_2'^+)]^l (i\hat{D}_3'^-)^p \varphi_0^{j,-}$$

$$D_1^{\prime \mp} = \partial_1 \mp M_1 \bar{z}^{\bar{1}}$$
$$D_2^{\prime \mp} \propto \partial_2 \pm \kappa \bar{z}^{\bar{2}}$$
$$D_3^{\prime \mp} \propto i \left(\partial_2 \mp \kappa \bar{z}^{\bar{2}}\right)$$

#### Warped zero modes (cont.)

Expand warped zero mode in terms of unwarped massive modes



• Examples given in [1011.xxxx]

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#### Adjoint fields

Recall the adjoint field wavefunctions

$$\psi_0 \sim e^{-3\alpha/2} \qquad \psi_{1,2} \sim e^{\alpha/2} \qquad \psi_3 \sim e^{-3\alpha/2}$$

•The resulting 4D kinetic terms are

#### Chiral matter fields

•Warping modifications for chiral matter more complex

$$S = -\frac{1}{g_8^2} \int_{\mathcal{W}} d^8 x \sqrt{\tilde{g}} \operatorname{tr} \left\{ \frac{1}{2} \eta^{\mu\nu} \tilde{g}^{ab} F_{\mu a} F_{\nu b} + e^{-4\alpha} \eta^{\mu\nu} \tilde{g}_{ij} \partial_\mu \Phi^i \partial_\nu \Phi^j \right\}$$
$$\downarrow$$
$$\mathcal{K}_{j\bar{k}}^{\mp} = \frac{1}{g_8^2 \mathcal{V}_{\mathbf{w}}} \int_{\mathcal{S}_4} d^4 y \left( \operatorname{Im} \tau \right)^{-1} \left( \mathbf{X}^{k,\mp} \right)^* \cdot e^{\#\alpha} \mathbf{X}^{j,\mp}$$
$$e^{\#\alpha} = \operatorname{diag} \left( e^{-4\alpha}, 1, 1, e^{-4\alpha} \right)$$

•  $\mathbf{X}^{j,\mp}$  is the warped zero mode, not simply related to unwarped zero mode

# Chiral matter fields (cont.)

In the weak warping limit, the first order correction is

$$\delta \mathcal{K}_{j\bar{k}}^{\mp} = \frac{1}{\mathcal{V}} \int_{\mathcal{S}_4} \mathrm{d}^4 y \left( \mathrm{Im}\,\tau \right)^{-1} \left( \chi_3^{k,\mp} \right)^* \beta \chi_3^{j,\mp} - \frac{\delta \mathcal{V}}{\mathcal{V}} \left( \mathcal{K}_{j\bar{k}}^{\mp} \right)_0$$
  
unwarped

 Generic warp factors will introduce off-diagonal terms in the Kähler metric

 Second order corrections make use of warped wavefunctions (work in progress...)

# Summary and future directions

- Studied the wavefunctions for bifundamental matter in warped compactifications in both the chiral and non-chiral cases
- Needed to develop warped effective field theory and detailed phenomenology (though still work to be done!)
- Warping effects are more intricate than for adjoint matter; require a series expansion in unwarped massive modes
- General warp factors induce off-diagonal terms in the Kähler metric
- Extension to Calabi-Yau case is likely tricky...
- With a non-SUSY source (such as D3-branes), wavefunctions can be used to study soft terms (work in progress)