

(1)

## Physics 131 Wi 2005 HW#5

ch 4.

Qu. 12.

$$a = \frac{v^2}{R} \quad \because R_3 > R_4 = R_1 > R_2 \Rightarrow a_3 < a_4 = a_1 < a_2$$

Pro. 44.

$$a). T = \left( \frac{1200 \text{ rev}}{\text{min}} \right)^{-1} = 8.3 \times 10^{-4} \text{ min} = 0.050 \text{ s}$$

$$C = 2\pi r = 2\pi(0.15) = 0.94 \text{ m}$$

$$b). v = \frac{C}{T} = \frac{0.94}{0.050} = 19 \text{ m/s}$$

$$c). a = \frac{v^2}{r} = \frac{19^2}{0.15} = 2.4 \times 10^3 \text{ m/s}^2$$

$$d). T = 0.050 \text{ s}$$

Pro. 45.

$$a). \text{ From } a = \frac{v^2}{r}, \text{ we have:}$$

$$v = \sqrt{ra} = \sqrt{(5.0 \text{ m})(7.0)(9.8 \text{ m/s}^2)} = 19 \text{ m/s}$$

$$b). \text{ The period: } T = \frac{2\pi r}{v} = 1.7 \text{ s}$$

In one minute (60s), the astronaut executes

$$\frac{t}{T} = \frac{60}{1.7} = 35 \text{ revolutions.}$$

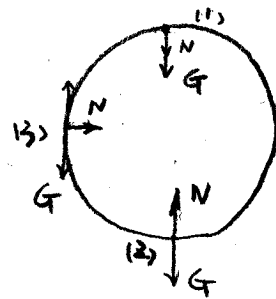
$$c). T = 1.7 \text{ s}$$

ch 6

Qu. 10

(a) From  $a = \frac{v^2}{R}$ ,  $v = \text{constant}$ ,

we have:  $a_{(1)} = a_{(2)} = a_{(3)}$



(b)  $F = ma$

$\Rightarrow F_{(1)} = F_{(2)} = F_{(3)}$  (net centripetal force)

(c)

For (1):  $F_{(1)} = N_{(1)} + G \Rightarrow N_{(1)} = F_{(1)} - G$

For (2):  $F_{(2)} = N_{(2)} - G \Rightarrow N_{(2)} = F_{(2)} + G$

For (3):  $F_{(3)} = N_{(3)}$

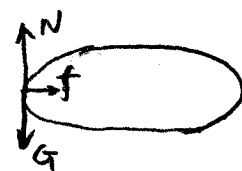
so, we have:  $N_{(2)} > N_{(3)} > N_{(1)}$

Pro. 39

If the bicycle does not slip,

$$f \leq f_{s, \text{max}} = \mu_s N = \mu_s mg,$$

and  $f = m \frac{v^2}{R}$



so, we have:

$$m \frac{v^2}{R} \leq \mu_s mg \Rightarrow R_{\text{min}} = \frac{v^2}{\mu_s g} = \frac{(29 \times \frac{1000 \text{ m}}{3600 \text{ s}})^2}{(0.32)(9.8 \text{ m/s}^2)} = 21 \text{ m}$$

Pro. 43

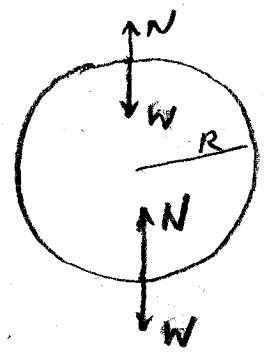
a) At the top,  $N < W$ , which means the seat is pushing up with a force that is smaller than the student's weight, and we say the student experiences a decrease in his "apparent weight" at the highest point.

(b)

At the top, we have:

$$W - N = \frac{mv^2}{R}$$

$$\Rightarrow \frac{mv^2}{R} = W - N = 667 - 556 = 111 \text{ N}$$



At the lowest point,

$$N - W = \frac{mv^2}{R}, \text{ because the Ferris wheel is "steadily rotating".}$$

$$\text{Then, } N = W + \frac{mv^2}{R} = 667 + 111 = 778 \text{ N}$$

(c) If the speed is doubled,

$$\frac{m(2v)^2}{R} = 4 \frac{mv^2}{R} = 4 \times 111 = 444 \text{ N}$$

At the highest point,

$$W - N = 444 \text{ (N)} \Rightarrow N = 667 - 444 = 223 \text{ N}$$

At the lowest point,

$$N - W = 444 \text{ (N)} \Rightarrow N = 667 + 444 = 1.1 \text{ kN}$$

**Pro. 44.**

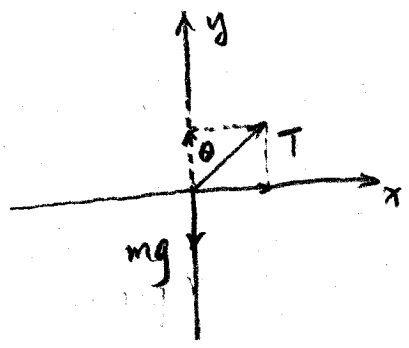
$$v = 16 \text{ km/h} = 4.4 \text{ m/s}$$

x-direction:

$$T \sin \theta = m \frac{v^2}{R} \quad (1)$$

y-direction:

$$T \cos \theta = mg \quad (2)$$



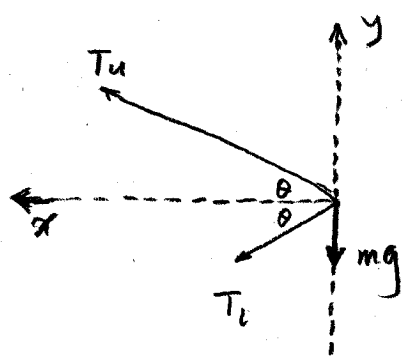
$$\text{Solving Eqs. (1) and (2), we get: } \theta = \tan^{-1} \left( \frac{v^2}{Rg} \right) = \tan^{-1} \left( \frac{4.4^2}{9.1 \times 9.8} \right) = 12^\circ$$

pro. 47.

a) Free-body diagram.

$T_u$ : tension exerted by the upper string

$T_l$ : tension exerted by the lower string



b) Take the +x direction to be leftward, +y upward.

x-direction:

$$T_u \cos \theta + T_l \cos \theta = \frac{mv^2}{R} \quad (1)$$

y-direction:

$$T_u \sin \theta - T_l \sin \theta - mg = 0 \quad (2)$$

Solving Eq. (2), we get

$$T_l = T_u - \frac{mg}{\sin \theta} = 35 - \frac{1.34 \times 9.8}{\sin 30^\circ} = 8.74 \text{ N}$$

c) The net force is leftward,

$$F_{\text{net}} = T_u \cos \theta + T_l \cos \theta = (T_u + T_l) \cos \theta = (35 + 8.74) \cos 30^\circ = 37.9 \text{ N}$$

d)

$$\tan \theta = \frac{(1.7)}{2} \Rightarrow R = \frac{(1.7)}{\tan 30^\circ} = 1.47 \text{ m}$$

$$F_{\text{net}} = m \frac{v^2}{R} \Rightarrow v = \sqrt{\frac{R F_{\text{net}}}{m}} = \sqrt{\frac{1.47 \times 37.9}{1.34}} = 6.45 \text{ m/s}$$

ch. 14.

pro. 16

a) The gravitational acceleration at the surface of the Moon.

$$g_{\text{moon}} = 1.67 \text{ m/s}^2 \quad (\text{see Appendix C})$$

and, we have:  $W_{\text{earth}} = m g_{\text{earth}}$

$$W_{\text{moon}} = m g_{\text{moon}}$$

$$\text{so, } \frac{W_{\text{earth}}}{W_{\text{moon}}} = \frac{g_{\text{earth}}}{g_{\text{moon}}} \Rightarrow W_{\text{moon}} = 100 \text{ N} \frac{1.67}{9.8} = 17 \text{ N}$$

b) For the force on that object caused by Earth's gravity to equal 17N, then the free-fall acceleration at its location must be  $a_g = 1.67 \text{ m/s}^2$ .

Thus,

$$a_g = \frac{GM_E}{r^2} \Rightarrow r = \sqrt{\frac{GM_E}{a_g}} = 1.5 \times 10^7 \text{ m}$$

So, the object would need to be a distance of  $\frac{r}{R_E} = 2.4$  "radii" from Earth's center.

Pro. 21

From Eq. 14-13, we see the extreme case is when "g" becomes zero, and plugging in Eq. 14-14 leads to

$$0 = \frac{GM}{R^2} - R\omega^2 \Rightarrow M = \frac{R^3\omega^2}{G}$$

Thus, with  $R = 20000 \text{ m}$  and  $\omega = 2\pi \text{ rad/s}$ ,

$$\text{we find } M = 4.7 \times 10^{24} \text{ kg.}$$