

## Photoelectric effect in hydrogen

Consider a hydrogen atom in its ground state  $|100\rangle$ , centered at the origin, exposed to an incident electromagnetic wave in Coulomb gauge ( $\vec{\nabla} \cdot \vec{A} = 0$ )

$$\vec{A}(\vec{r}, t) = \vec{A}_0 \cos(\vec{k} \cdot \vec{r} - \omega t).$$

If  $\omega$  is large enough, this can ionize the atom.

Let's try to calculate the ionization rate using

Fermi's golden rule:

$$R_{i \rightarrow f} \equiv \frac{dP_{i \rightarrow f}}{dt} = \frac{2\pi}{\hbar} |\langle f^0 | \hat{H}' | i^0 \rangle|^2 \delta(E_f^0 - E_i^0 - \hbar\omega)$$

We don't need the term  $\sim \delta(E_f^0 - E_i^0 + \hbar\omega)$  arising from the  $e^{i\omega t}$  part of the cosine because we are already in the lowest state of the atom.

The final state is a positive energy eigenstate in the continuum. We should use a Coulomb wave (i.e. a positive energy ( $E > 0$ ) solution of the Coulomb Hamiltonian). It turns out that, for subtle reasons that we don't have time to explore in detail (see Bethe and Salpeter, *The Theory of One- and Two-electron Atoms*, Plenum 1977), for excitation from s-states (but not from p-states!) we can replace the Coulomb wave by a plane wave,  $|p_f\rangle \leftrightarrow \frac{e^{-i\vec{p}_f \cdot \vec{r}}}{(2\pi\hbar)^{3/2}}$ .

We write the interaction Hamiltonian in Coulomb gauge as (remember  $\vec{p} \rightarrow \vec{p} - \frac{q}{c} \vec{A}$ )

$$\hat{H}'(t) = -\left(\frac{e}{c}\right) \frac{1}{2m} (\vec{A} \cdot \hat{\vec{p}} + \hat{\vec{p}} \cdot \vec{A}) \quad (\text{ignore term } \sim |\vec{A}|^2 \text{ which is second order in } \vec{A})$$

$$(\vec{\nabla} \cdot \vec{A} = 0)$$

$$= \frac{e}{mc} \vec{A} \cdot \hat{\vec{p}} = \frac{e}{mc} \cos(\vec{k} \cdot \vec{r} - \omega t) \vec{A}_0 \cdot \hat{\vec{p}}$$

$$= \frac{e}{2mc} \left[ e^{i(\vec{k} \cdot \vec{r} - \omega t)} + e^{-i(\vec{k} \cdot \vec{r} - \omega t)} \right] \vec{A}_0 \cdot \hat{\vec{p}}$$

(ignore term  $\sim e^{i\omega t}$ )

$$\rightarrow \frac{e}{2mc} e^{i\vec{k} \cdot \vec{r}} \vec{A}_0 \cdot \hat{\vec{p}} e^{-i\omega t} \equiv \hat{H}'(\vec{r}) e^{-i\omega t}$$

$$\Rightarrow \langle f^0 | \hat{H}' | i^0 \rangle = \frac{e}{2mc} \frac{1}{(2\pi\hbar)^{3/2}} \sqrt{\frac{1}{\pi a_0^3}} \int d^3r e^{-i\vec{p}_f \cdot \vec{r} / \hbar} e^{i\vec{k} \cdot \vec{r}} e^{-r/a_0} \times \vec{A}_0 \cdot (-i\hbar \vec{\nabla}) e^{-r/a_0}$$

The factor  $e^{i\vec{k} \cdot \vec{r}}$  adds a momentum  $\hbar\vec{k}$  to the atom  $\rightarrow$  the electromagnetic wave imparts momentum on the atom. How does this affect the energetics?

To ionize the atom, we must transfer an energy of

$$\text{order } 1R_y: \quad \hbar\omega \sim \frac{e^2}{a_0} \Rightarrow \hbar k = \hbar \frac{\omega}{c} \sim \frac{e^2}{a_0 c}$$

In the ground state, the electron has a typical momentum (from the uncertainty relation)

$$p \sim \frac{\hbar}{a_0}$$

$$\Rightarrow \frac{\hbar k}{p} \approx \frac{e^2}{a_0 c} \cdot \frac{a_0}{\hbar} = \frac{e^2}{\hbar c} = \alpha = \frac{1}{137}$$

So  $\frac{\hbar k}{p} \ll 1$  unless we give the final state electron an energy of  $O(137 \text{ Ry})$ . So let's work in this energy domain where we can use  $\hbar k \ll p$ .

In principle we should also account for the interaction of the electron spin with the  $\vec{B}$ -field associated with the incident  $\vec{A}(\vec{r}, t)$ :

$$\frac{\left\langle \frac{e}{2mc} \vec{S} \cdot \vec{B} \right\rangle}{\left\langle \frac{e}{mc} \vec{A} \cdot \vec{p} \right\rangle} \approx \frac{\left\langle \hbar \vec{\sigma} \cdot (\nabla \times \vec{A}) \right\rangle}{4 \left\langle \vec{A} \cdot \vec{p} \right\rangle} \approx \frac{1}{4} \frac{\hbar k}{p} \ll 1$$

So we can ignore the  $\vec{\mu}_s \cdot \vec{B}$  interaction for the same reason.

$$\text{Since } \hbar k \ll p \sim \frac{\hbar}{a_0} \Rightarrow \boxed{ka_0 \ll 1}$$

So for ionization into final states with  $E_f \sim O(\text{few Ry})$ , the wavelength of the light is much larger than

the Bohr radius of the 1s electron

→  $\vec{A}$  is basically  $\vec{r}$ -independent over the range of the atom in its initial state.

→ we may approximate  $e^{i\vec{k}\cdot\vec{r}} \approx 1$  under the integral ("dipole approximation")

The atom sees effectively a spatially constant electric field whose magnitude oscillates in time with frequency  $\omega$ :

$$\vec{E} = -\frac{1}{c} \frac{\partial \vec{A}}{\partial t} = -\frac{1}{c} \frac{\partial}{\partial t} \left( \frac{\vec{A}_0}{2} e^{-i\omega t} \right) = \frac{i\omega}{2c} \vec{A}_0 e^{-i\omega t}$$

Now let's work out the matrix element in dipole approximation:

Let's integrate first by parts:

$$\int d^3r e^{-i\vec{p}_f \cdot \vec{r}/\hbar} \vec{A}_0 \cdot (-i\hbar \vec{\nabla}) e^{-r/a_0} = \int d^3r (i\hbar \vec{\nabla} e^{-i\vec{p}_f \cdot \vec{r}/\hbar}) \cdot \vec{A}_0 e^{-r/a_0}$$

$$= \vec{p}_f \cdot \vec{A}_0 \int d^3r e^{-i\vec{p}_f \cdot \vec{r}/\hbar} e^{-r/a_0} =$$

$$= \vec{p}_f \cdot \vec{A}_0 \int_0^{2\pi} d\varphi \int_0^\infty r^2 dr e^{-r/a_0} \underbrace{\int_{-1}^1 d(\cos\theta) e^{-ip_f r \cos\theta/\hbar}}_{\frac{e^{-ip_f r/\hbar} - e^{ip_f r/\hbar}}{-ip_f r/\hbar}}$$

$$= 2\pi \hbar \frac{\vec{p}_f \cdot \vec{A}_0}{p_f} i \left( -\frac{\partial}{\partial (1/a_0)} \right) \int_0^\infty dr \left( e^{-\left(\frac{1}{a_0} + i\frac{p_f}{\hbar}\right)r} - e^{-\left(\frac{1}{a_0} - i\frac{p_f}{\hbar}\right)r} \right)$$

$$= \frac{8\pi/a_0}{\left[ \frac{1}{a_0^2} + \left(\frac{p_f}{\hbar}\right)^2 \right]^2} \vec{p}_f \cdot \vec{A}_0$$

Putting this together with the normalization factors and squaring it we finally get

$$R_{i \rightarrow f} = \frac{2\pi}{\hbar} \left(\frac{e}{2mc}\right)^2 \frac{1}{(2\pi\hbar)^3} \frac{1}{\pi a_0^3} \frac{|\vec{A}_0 \cdot \vec{p}_f|^2}{\left[1 + \left(\frac{p_f a_0}{\hbar}\right)^2\right]^4} 64\pi^2 a_0^6 \delta(E_f^0 - E_i^0 - \hbar\omega)$$

Experimentally, we can't measure  $E_f$  with infinite accuracy, so we must integrate this rate over the finite energy resolution. This means we are only interested in the area under the  $\delta$ -function which is peaked at

$$\frac{p_f^2}{2m} = E_i^0 + \hbar\omega \quad \rightarrow \quad p_f = \sqrt{2m(E_i^0 + \hbar\omega)}$$

So let's count the number of electrons per unit time emitted into solid angle  $d\Omega$ :

$$\begin{aligned} R_{i \rightarrow d\Omega} &= \int_{p_f - \epsilon}^{p_f + \epsilon} p^2 dp d\Omega \frac{4e^2 a_0^3}{m^2 \hbar^4 c^2} \frac{\vec{A}_0 \cdot \vec{p}_f}{\left[1 + \left(\frac{p_f a_0}{\hbar}\right)^2\right]^4} \underbrace{\delta\left(\frac{p^2}{2m} - \frac{p_f^2}{2m}\right)}_{\frac{m}{p_f} \delta(p - p_f)} \\ &= \frac{4e^2 a_0^3}{m^2 \hbar^4} \frac{p_f |\vec{p}_f \cdot \vec{A}_0|^2}{\left[1 + \left(\frac{p_f a_0}{\hbar}\right)^2\right]^4} d\Omega \end{aligned}$$

The  $\vec{p}_f \cdot \vec{A}_0$  factor tells us that the emission is biased towards the direction of  $\vec{A}_0$  which is also the direction of the electric field of the light that rips out the electron from its bound state.

If we had kept the  $e^{i\vec{k}\cdot\vec{r}}$  factor, we would see that there is an additional (weaker) bias towards  $\vec{k}$ , reflecting the momentum transfer from the light wave.

(We can actually do this calculation exactly, by simply combining  $e^{-i\vec{p}_f\cdot\vec{r}/\hbar}$   $e^{i\vec{k}\cdot\vec{r}}$  under the integral and replacing  $\vec{p}_f \rightarrow \vec{p} - \hbar\vec{k}$  in the result:

$$R \sim |\vec{A}_0 \cdot (\vec{p}_f - \hbar\vec{k})|^2$$

This is largest when  $\vec{p}_f \parallel \vec{k}$  and  $\vec{p}_f \parallel \vec{E} \sim \vec{A}_0$ .)

The total ionization rate is obtained by integrating over the solid angle:  $\int_{-1}^1 \cos^2\theta \, d\cos\theta = \frac{2}{3}$ ,  $\int_0^{2\pi} d\phi = 2\pi$

$$R_{i \rightarrow \text{all}} = \frac{16 a_0^3 e^2 p_f^3 |\vec{A}_0|^2}{3 m c^2 \hbar^4 (1 + (p_f a_0 / \hbar)^2)^4}$$

total ionization rate

The amount of energy absorbed from the incoming light per unit time is

$$\frac{dE_{\text{abs}}}{dt} = \hbar\omega R_{i \rightarrow \text{all}}$$

How does this compare with the energy contained in the incoming light wave?

The incoming plane wave brings in energy at

$$\text{a rate of } |\vec{S}| = \frac{\omega^2 |\vec{A}_0|^2}{4\pi c} \langle \sin^2(\vec{k}\cdot\vec{r} - \omega t) \rangle_t = \frac{\omega^2 |\vec{A}_0|^2}{8\pi c} \text{ per unit area.}$$

If we put a perfectly black (100% absorbing) disk of area  $\sigma$  into that beam, it absorbs energy at the rate

$$\frac{\omega^2 |\vec{A}_0|^2}{8\pi c} \cdot \sigma = \frac{dE_{\text{abs}}}{dt}$$

We can compare this with the energy absorbed by the hydrogen atom and assign to it an effective black disk area, called the "photoelectric cross section" (or ionization cross section)

$$\sigma_{\text{ioniz}} = \frac{8\pi c}{\omega^2 |\vec{A}_0|^2} \cdot \hbar \omega \cdot R_{i \rightarrow \text{all}} = \frac{128 a_0^3 \pi e^2 p_f^3}{3 m c \omega \hbar^3 [1 + (\frac{p_f a_0}{\hbar})^2]^4}$$

The differential cross section  $\frac{d\sigma}{d\Omega}$  is defined

correspondingly as

$$\frac{d\sigma}{d\Omega} = \frac{8\pi c}{\omega^2 |\vec{A}_0|^2} \hbar \omega \frac{dR_{i \rightarrow d\Omega}}{d\Omega} = \frac{32 a_0^3 e^2 p_f^3 \cos^2 \theta}{m c \omega \hbar^3 [1 + (\frac{p_f a_0}{\hbar})^2]^4}$$

For  $\frac{p_f a_0}{\hbar} \gg 1$  we can ignore the 1 in the denominator:

$$\begin{aligned} \frac{d\sigma}{d\Omega} &\xrightarrow{p_f a_0 \gg \hbar} \frac{32 e^2 \hbar^5 \cos^2 \theta}{m c \omega p_f^5 a_0^5} && (\text{remember: } p_f a_0 \approx \hbar) \\ &= 32 \frac{\alpha^2 a_0}{\hbar \omega} \frac{\hbar c}{(p_f a_0 / \hbar)^5} \end{aligned}$$

(units of  $\sigma$  are barn =  $10^{-28} \text{m}^2$  or mb =  $0.1 \text{fm}^2$ )

## Spontaneous radiative decay of excited atoms

This requires treating the electromagnetic field as a quantum system ("field quantisation"), leading into quantum field theory. I will leave this for an advanced quantum theory course.