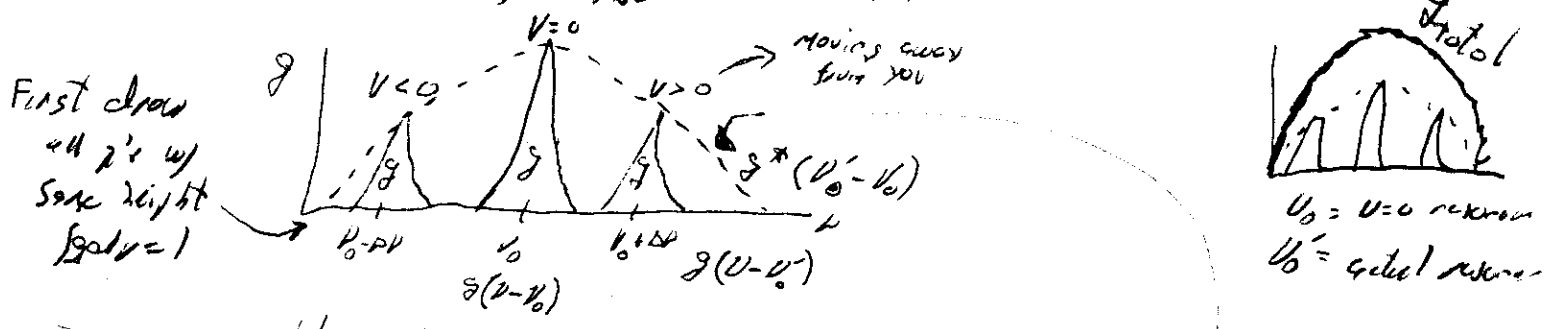


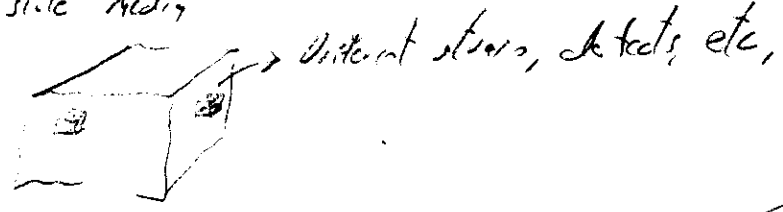
Now, for the inhomogeneous case.

We imagine the total population can be divided into sub-populations, each of a homogeneous line shape.

Eg: Atomic gas. The atoms are in motion and their emitted radiation is Doppler shifted and the radiation they "observe" is shifted as well.



Eg: Solid state media



different atoms, defects, etc,

atoms/volume

$$\equiv \int N_t \equiv N_t g^*(\nu'_0 - \nu_0) \Delta \nu'_0$$

$$\int \omega_t = N_t$$

with resonance frequency between ν'_0 & $\nu_0 + \Delta \nu'_0$ (but w/ a standard line shape about ν_0)

We had $\frac{dP}{d\nu} = N_t \omega_h(\nu - \nu_0) h\nu$ ↓ constant

↑ above at ν'_0

↑ resonance substance

↑ light photon - whatever color we want it to be

$$\int \left(\frac{dP}{d\nu} \right) = \int N_t \omega_h(\nu - \nu'_0) h\nu$$

$$= \left[N_t g^*(\nu'_0 - \nu_0) \Delta \nu'_0 \right] \omega_h(\nu - \nu_0) h\nu$$

power absorbed just = by the $\int N_t$ per unit volume

$$\frac{dP}{d\nu} = N_t h\nu \int \omega_h(\nu - \nu'_0) g^*(\nu'_0 - \nu_0) \Delta \nu'_0$$

↘ response of a population of atoms w/ resonance freq ν_0

↗ \propto # atoms/volume w/ resonance freq ν_0

↘ response of a population of atoms w/ resonance freq ν_0

For the homogeneous case we had.

$$\frac{dP}{dV} = N_A h \nu W_h$$

So we write for the inhomogeneous case:

$$\frac{dP}{d\nu} = N_A h \nu W_{in}$$

$$W_{in} = \int W_h(\nu - \nu_0') \gamma^* (\nu_0' - \nu_0) d\nu_0'$$

Smearing or broadening of homogeneous line

If we define:

$$G_{in} = \frac{W_{in}}{F}$$

we get

$$G_{in} = \int G_h(\nu - \nu_0') \gamma^* (\nu_0' - \nu_0) d\nu_0'$$

This is an average cross-section that we can use to characterize the whole system! (For some problems.)

Note: $G_h = G_h(\nu) \propto \nu \gamma^* (\nu - \nu_0)$
 $G_{in} = G_{in}(\nu) \propto \int \dots d\nu_0'$

Recall, we had:

$$\sigma_A = \frac{2\pi^2}{3n\epsilon_0 ch} |\mu|^2 \nu g(\nu - \nu_0)$$

→ atomic line shape

We can write:

$$\sigma_{in} = \frac{2\pi^2}{3n\epsilon_0 ch} |\mu|^2 \nu g_{in}(\nu - \nu_0)$$

→ total line shape

$$\sigma_{in} = \frac{W_{in}}{F} \dots$$

$$g_{in} = \int_{-\infty}^{\infty} g^*(x) g(\nu - \nu_0 - x) dx$$

x = change in resonance freq
x = $\nu'_0 - \nu_0$

limiting case #1) No inhomogeneous broadening

$$g^* = \delta(x)$$

$$g_{in} = g(\nu - \nu_0)$$

#2) Narrow atomic line

$$g = \delta(\nu - \nu_0 - x)$$

$$g_{in} = g^*(\nu - \nu_0)$$

e.g. Doppler shift provides all the broadening

Following the text, from now on, we'll just use:

$$\sigma = \frac{2\pi^2}{3n\epsilon_0 ch} |\mu|^2 \nu g(\nu - \nu_0)$$
$$W = \sigma F$$

and not specify broad or inhom. broadening explicitly. That's all in $g(\nu)$.

So, how does general 2-level system affect the light?

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We have $dF = -\sigma N_1 F dz$

All atoms in lower state,
 $N_1 = \# \text{ density}$

$dF = +\sigma N_2 F dz$

All atoms in upper state
General
 N_1 is lower state
 N_2 is upper

$dF = -\sigma(N_1 - N_2) F dz$

$dF = dF_{abs} + dF_{em}$

Assume $N_1 > N_2$

$\alpha \equiv \sigma(N_1 - N_2) = \text{absorption coefficient} \geq 0$

$\frac{dF}{dz} = -\alpha F$

$F(z) = F_0 e^{-\alpha z}$



$\alpha = \frac{1}{\text{length}}$

$\alpha = 2.303$ (or 2.3) (constant slope used)
 means flux falls by 1/e in 2.3m

But $N_2 > N_1$ (population inversion)

$\alpha \equiv \sigma(N_2 - N_1) \geq 0$

$\frac{dF}{dz} = \alpha F$

$F(z) = F_0 e^{\alpha z}$



Of course, once F becomes large enough, its absorption/emission affects $(N_1 - N_2)$.
 It drives it to zero. No absorption or gain.
 If $F=0$ $F(z)=0$.

Einstein's thermodynamic treatment (Non-degenerate treatment) How are stim em, spont em & abs related? (2)

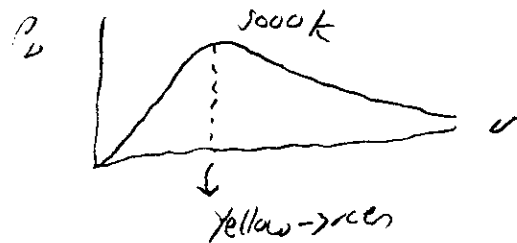
Even w/o a quantum treatment of light, we can use nature and a little physics to say some things about spont. emissio..

Suppose we have a black body at temp T .
By def'n, it is at equilibrium.
(We'll wait until it is.)

We know (Eq 2.2.22)

energy density $\rho(\nu) = \frac{8\pi\nu^3}{c^3} \frac{h\nu}{e^{h\nu/kT} - 1}$

DO NOT ERASE



Now, we have abs, stim em, spont em taking place in equilibrium
Characterize the rates as: $W_{21}^{sp} \propto I \propto \rho$

$W_{21}^{st} = B_{21} \rho_{\nu_0}$ stim em. $\rho = \rho(\nu_0)$

$W_{12}^{ab} = B_{12} \rho_{\nu_0}$ abs

$W_{21}^{sp} = A$ spont-em

B_{12}, B_{21}, A contain all atomic information now.
We require:

energy density is more convenient for a completely closed cavity. light's not going anywhere.

$B_{21} \rho_{\nu_0} N_2 + A N_2 = B_{12} \rho_{\nu_0} N_1$

$\rho \propto N_1 + N_2$
 ρ_{ν_0} independent of time.

However, from the Boltzmann distribution:

$\frac{N_2}{N_1} = e^{-h\nu_0/kT}$

(Non-degenerate case. Extension to degenerate case not interesting for us.)

$$B_{21} P_{\nu_0} + A N_2 = \frac{B_{12} P_{\nu_0}}{e^{-h\nu_0/kT}}$$

$$P_{\nu_0} = \frac{A}{B_{12} e^{h\nu_0/kT} - B_{21}} = \frac{A/B_{21}}{\frac{B_{12}}{B_{21}} e^{h\nu_0/kT} - 1}$$

Compare to Planck's law:

$$\frac{B_{12}}{B_{21}} = 1 \quad \Rightarrow \quad B_{12} = B_{21} = B \quad \checkmark$$

$$\frac{A}{B} = \frac{8\pi h \nu_0^3 n^3}{c^3}$$

Text shows,

$$B \propto |\mu|^2$$

$$\text{so } A \propto |\mu|^2$$

A strong transition will suffer large spont. emission losses.

$$P(\omega) = \frac{8\pi\nu^3}{c^3} \frac{h\nu}{e^{h\nu/kT} - 1}$$

$$P = \frac{A/B_{12}}{\frac{2}{B_{12}} e^{h\nu/kT} - 1}$$

$$\frac{B_{12}}{B_{21}} = 1 \quad B_{12} = B_{21} = B$$

$$\frac{A}{B} = \frac{8\pi h \nu^3}{c^3}$$

Text then goes on to show:

$$A = \frac{16\pi^3 \nu^3 \hbar^2}{3\hbar \epsilon_0 c^3}$$

$$B = \frac{2\pi^2 \hbar^2}{3\hbar^2 \epsilon_0 \hbar^2}$$

→ We have been interested in EM wave interaction

W_{12}	W_{21}	by calculation	↪	$\sigma \rightarrow \mu, z$
A		by formal	↪	$\sigma \rightarrow \mu, y$

Next → look at forms of S

Then → remove some of our approximations: 2-level system