

Black-Body Radiation

matter at temperature T
radiates electromagnetic waves

black-body: absorbs all incident radiation
reflects none
emits smooth spectrum

Stefan's Law 1879

power P radiated per area A
depends only on temperature T

$$P = \sigma A T^4$$

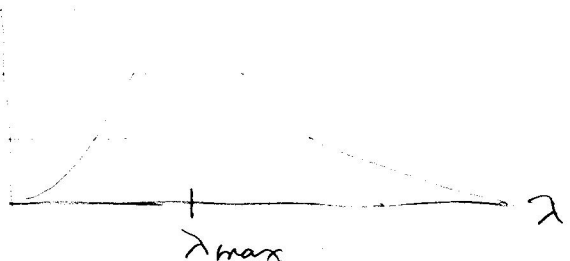
$$\sigma \approx 6 \times 10^{-8} \frac{\text{J}}{\text{A} \cdot \text{m}^2 \cdot \text{K}^4} \quad \text{universal constant}$$

Wien's Law 1893

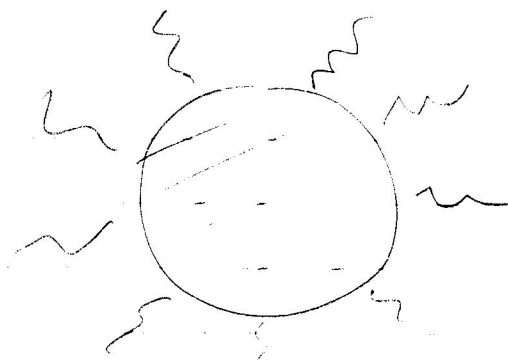
blackbody spectrum has a maximum
at a wavelength λ_{max} that depend only on T

$$\lambda_{\text{max}} = \frac{w}{T}$$

$$w \approx 3 \cdot 10^{-3} \text{ m} \cdot \text{K}$$

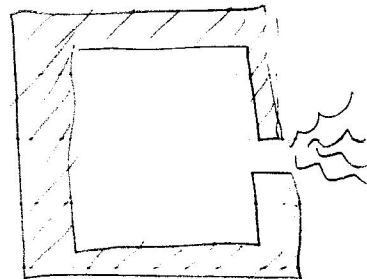


Radiation from black body at temperature T



is identical to

Radiation emitted by hole in cavity whose walls are at temperature T



$$P_{\text{hole}} = \sigma A_{\text{hole}} T^4$$

Travelling EM waves emitted from hole are related to standing EM waves in cavity

$$P_{\text{hole}} = A_{\text{hole}} \left(\frac{1}{4} c \right) \rho_{\text{cavity}}$$

↑
angle average of velocity of light

ρ_{cavity} = energy density of standing EM waves

ρ_{cavity} can be expressed

as integral over frequency of normal modes

$$\rho_{\text{cavity}} = \int_0^{\infty} \frac{8\pi\nu^2 d\nu}{c^3} \bar{E}(\nu)$$



sum over normal modes
with frequency ν
per volume

average energy
of normal mode
of frequency ν .

frequency spectrum of black-body radiation

$$dP = A \left(\frac{1}{4c} \right) \frac{8\pi}{c^3} \bar{E}(\nu) \nu^2 d\nu$$

$\bar{E}(\nu)$ = average energy
in standing EM wave with frequency ν
in equilibrium at temperature T

Classical thermodynamics

harmonic oscillator

in equilibrium at temperature T

has energy distribution $\propto e^{-E/kT}$

average energy

$$\begin{aligned} \bar{E} &= \frac{\int_0^{\infty} dE E e^{-E/kT}}{\int_0^{\infty} dE e^{-E/kT}} \\ &= \frac{(kT)^2}{kT} = kT \end{aligned}$$

predicted black-body spectrum

$$dP = A \frac{2\pi}{c^3} kT \nu^2 d\nu$$

Planck: fitting formula for spectrum

$$dP = A \cdot \frac{2\pi}{c^2} kT \frac{h\nu/kT}{e^{h\nu/kT} - 1} \nu^2 d\nu$$

where h is a universal constant

$$h \approx 7 \times 10^{-34} \text{ J}\cdot\text{s}$$

Suppose energy of normal mode of frequency ν can only be integer multiple of $h\nu$

$$E = 0 \text{ or } h\nu \text{ or } 2h\nu \text{ or } \dots$$

$$E = nhf, \quad n = 0, 1, 2, \dots$$

average energy?

$$\bar{E}(\nu) = \frac{\sum_{n=0}^{\infty} nh\nu e^{-nh\nu/kT}}{\sum_{n=0}^{\infty} e^{-nh\nu/kT}}$$

$$= \frac{kT^2 \frac{d}{dT} \frac{-1}{1 - e^{-h\nu/kT}}}{\frac{1}{1 - e^{-h\nu/kT}}}$$

$$= \frac{h\nu}{e^{h\nu/kT} - 1}$$

integrate Planck's formula over ν
to get total power

$$P = A \frac{2\pi h}{c^2} \int_0^{\infty} \frac{\nu^3 d\nu}{e^{h\nu/kT} - 1}$$

change to dimensionless variable

$$x = \frac{h\nu}{kT} \quad dx = \frac{h}{kT} d\nu$$

$$\nu^3 d\nu = \left(\frac{kT}{h}\right)^4 x^3 dx$$

$$P = A \frac{2\pi h}{c^2} \left(\frac{kT}{h}\right)^4 \int_0^{\infty} \frac{x^3 dx}{e^x - 1}$$

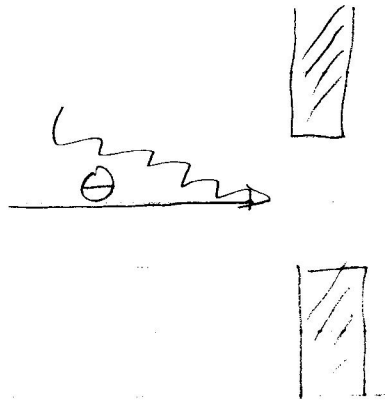
$$= A \sigma T^4$$

Stefan's constant

$$\sigma = \frac{2\pi \cdot k^4}{c^2 h^3} \underbrace{\int_0^{\infty} \frac{x^3 dx}{e^x - 1}}_{0.49}$$

expressed in terms of fundamental constants
 c, k, h

Light waves near hole in cavity



velocity component of light⁺
perpendicular to hole: $v_x = c \cos \theta$

$0 < \theta < \frac{\pi}{2}$: $v_x > 0$ (approaching hole)

$\frac{\pi}{2} < \theta < \pi$: $v_x < 0$

angle average of v_x for light approaching hole

$$\frac{\int_0^{\frac{\pi}{2}} \sin \theta d\theta (c \cos \theta) + \int_{\frac{\pi}{2}}^{\pi} \sin \theta d\theta \cdot 0}{\int_0^{\pi} \sin \theta d\theta \cdot 1}$$
$$= \frac{c \int_0^{\frac{\pi}{2}} d\theta \sin \theta \cos \theta}{\int_0^{\pi} d\theta \sin \theta} = c \frac{\int_0^{\frac{\pi}{2}} d(\cos^2 \theta / 2)}{\int_0^{\pi} d(\cos \theta)}$$
$$= c \frac{1/2}{2} = \frac{c}{4}$$

Wave equation in 1 dimension
for waves with speed c


$$\left[\left(\frac{\partial}{\partial t} \right)^2 - c^2 \left(\frac{\partial}{\partial x} \right)^2 \right] f = 0$$

travelling waves: 

$$f(x, t) = \sin(kx - \omega t) \quad \omega = kc$$

standing waves in interval $0 < x < L$

that vanish at endpoints: $f = 0$ at $x = 0$
at $x = L$

$$f(x, t) = \sin(kx) \sin(\omega t) \quad \text{$$

$$k = \frac{n\pi}{L}, \quad n = 1, 2, 3, \dots \quad \omega = kc = \frac{n\pi c}{L}$$

discrete frequencies: $\nu_n = \frac{\omega_n}{2\pi} = \frac{c}{2L} n$

difference between adjacent frequencies

$$\Delta \nu_n = \frac{c}{2L} \Delta n, \quad \Delta n = 1$$

Sum over normal modes:

$$\sum_{n=1}^{\infty} 1 = \sum_{n=1}^{\infty} \Delta n = \sum_{n=1}^{\infty} \frac{2L}{c} \Delta \nu_n \rightarrow \frac{2L}{c} \int_0^{\infty} d\nu$$

Standing EM waves in 3 dimensions
in cubic cavity of length L

$$\vec{E}(x, y, z, t) = \vec{E}_0 \sin(k_1 x) \sin(k_2 y) \sin(k_3 z) \sin(\omega t)$$

2 polarizations \Rightarrow 2 solutions for \vec{E}_0

wave equation: $\omega^2 = c^2(k_1^2 + k_2^2 + k_3^2)$

boundary condition: $k_1 = \frac{n_1 \pi}{L}$ $n_1 = 1, 2, 3, \dots$
 $k_2 = n_2 \pi / L$ $n_2 = 1, 2, 3, \dots$
 $k_3 = n_3 \pi / L$ $n_3 = 1, 2, 3, \dots$

Sum over normal modes

$$\begin{aligned} 2 \sum_{n_1=1}^{\infty} \sum_{n_2=1}^{\infty} \sum_{n_3=1}^{\infty} &\rightarrow 2 \frac{L}{\pi} \int_0^{\infty} dk_1, \frac{L}{\pi} \int_0^{\infty} dk_2, \frac{L}{\pi} \int_0^{\infty} dk_3 \\ &= 2 \frac{L^3}{\pi^3} \int_{k_1>0, k_2>0, k_3>0} d^3 k \\ &= 2 \frac{L^3}{\pi^3} \frac{1}{8} 4\pi \int_0^{\infty} k^2 dk \quad k = \sqrt{k_1^2 + k_2^2 + k_3^2} \\ &= 2 \frac{L^3}{\pi^3} \frac{\pi}{2} \left(\frac{2\pi}{c}\right)^3 \int_0^{\infty} v^2 dv \quad = \frac{\omega}{c} \\ &= 2 \frac{4\pi}{c^3} L^3 \int_0^{\infty} v^2 dv \quad = \frac{2\pi v}{c} \\ &= V \cdot \frac{8\pi}{c^3} \int_0^{\infty} v^2 dv \end{aligned}$$